

Report from the Noble Liquids Group:

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Brief Introduction

- Dark Matter Detection, $0\nu 2\beta$, Neutrino Detectors
 - “Conventional” DM Detectors T. Shutt
 - Light Noble Detectors D. McKinsey
 - Directional Detectors M. Leyton
 - LAr TPC A. Rubbia
 - LAr Test Facilities at FNAL & CERN J. Raaf
 - $0\nu 2\beta$ with gas Xe, exotic $0\nu 2\beta$ D. Nygren
 - LAr R&D Program at BNL Y. Li
 - $0\nu 2\beta$ with liquid Xe G. Gratta
 - $0\nu 2\beta$ inXe with Ba tagging W. Fairbank

Findings

- LXe/LAr Dark Matter Detectors (Shutt)
 - Single or Two Phase Detectors, Ar or Xe
 - Scalable to O(10-100) tonnes
 - Two phase detectors use charge/scintillation for NR/ER discrimination
 - Single phase detectors rely on Pulse Shape Discrimination in Ar
 - High discrimination ~99.9 desirable for 30-50 tonne for LCT
 - HV continues to be a challenge
 - PTFE is “unreasonably reflective” in LXe
 - Scintillation light in Lar needs WLS.
 - ^{39}Ar background requires O(10^8) reduction
 - PSD >108 at thresholds 40-50 keV
 - Advent of low background PMTs has enabled current experiments

Findings

- Light Nobles (McKinsey)
 - He4 kinematic matching pulls up energy deposition for low mass DM, in window of masses between few hundred MeV - few GeV
 - Measurement of multiple signals - singlet, triplet scintillation, ionization, and heat could improve ER discrimination
 - LXe detectors doped with Ne could enhance sensitivity to low mass WIMPs
- Directional DM (Leyton)
 - Track shape and head/tail imaged in gas at low pressure ~ 0.5 Torr
 - WIMPs have preferred direction in galactic coordinates.
 - CYGNUS is confederation of the DDM community, mostly gas TPC's, emulsion. Some NITPC's. All small.

Findings

- LAR TPC (Rubbia)
 - Large international effort on a number of large scale prototype LAr TPCs.
 - Existing prototypes demonstrate bubble chamber-like tracks.
 - Plan to build 4 x 10 kT LAr modules at SURF. Membrane cryostats, long drift lengths.
- LAr Test Facilities (Raaf)
 - Extensive set of complementary test facilities exist at CERN and FNAL

Findings

- NEXT (high pressure Xe gas) (Nygren)
 - Excellent energy resolution, maybe $<5 \times 10^{-3}$ FWHM at 2457 KeV
 - Topology of events should be clear, with several cm tracks and ionization “meatballs” from stopping electrons. This provides additional discrimination of backgrounds.
 - Non-cryogenic, gas with density $<0.55 \text{ g/cm}^3$

Findings

- LAr R&D at BNL (Li)
 - Studies of fundamental properties of LAr, complementary to work at FNAL
- $0\nu 2\beta$ (Gratta, Fairbank)
 - EXO-200 discovered $2\nu 2\beta$ in ^{136}Xe .
 - Competitive limits on $0\nu 2\beta$ in ^{136}Xe .
 - nEXO is foreseen at 5 tonnes, which would be many γ interaction lengths, single drift space, SiPM photodetectors, and tile charge collection system.
 - Ba tagging would eliminate all backgrounds save $2\nu 2\beta$, but extremely non-trivial!

Comments

- Dark Matter Detectors
 - Enhanced light collection would enhance the reach of both detectors.
 - Directional detectors might ally with non-directional.
 - Directional to confirm detection likely 100 times larger in mass. Using directionality to push past neutrino floor would require kton or substantially larger mass
 - Underground Ar 1000-fold reduced in ^{39}Ar has been demonstrated, which should allow ~ 100 ton scale TPCs without strong pileup
- LAr TPC's
 - LAr TPCs are the necessary technology for DUNE, comprising LBNF, nucleon decay, and supernova neutrino detection.
 - Large scale LAr TPC's may be a new instrument for discovery.
 - First module will be single phase; two phase detectors under development.
 - Magnetized modules would be desirable, but may not be practical.

Comments

- $0\nu 2\beta$
 - High Xe purity is required for energy resolution
 - Tracking required for background rejection.
 - EXO-200 backgrounds are accurately modeled
 - nEXO HV may be difficult.
 - Ba tagging R&D in next 3 years is critical for an optimal low background nEXO.
- NEXT
 - Interesting idea for identifying Ba^{++} using single molecule fluorescence. Needs lots of work.

Identification of Risks and Opportunities

- Risks
 - Lack of sufficiently stable agency support
 - HV management for large liquid detectors
 - Inability to achieve sufficiently low backgrounds for DM, $\beta\beta$ decay
 - DUNE: Long drift requires ultra-high purity
- Opportunities
 - Increased science reach through improved light sensitivity and/or development of ultra-low background light readout.

Recommendations

- Increase funding(!)
- Areas that may need more support:
 - HV management in large detectors
 - Scintillation light detectors: DUNE, DM, nEXO (any synergies?)
 - Noble liquid purification in liquid state

Possible Grand Challenge Ideas

- Feasible high mass directional DM detector.
- Ba tagging in ^{136}Xe $0\nu 2\beta$
- Ultra-low background, high QE ($>40\%$), low dark rate, large area photosensor.